

---

# The X/Ka 2023b Celestial Frame

Christopher Jacobs<sup>\*1</sup>, Shinji Horiuchi<sup>2</sup>, Daniel Firre<sup>3</sup>, Yasuhiro Murata<sup>4</sup>, Hiroshi Takeuchi<sup>4</sup>, and Takahashi Uchimura<sup>4</sup>

<sup>1</sup>Jet Propulsion Laboratory - California Institute of Technology (JPL) – Jet Propulsion Laboratory  
4800 Oak Grove Drive Pasadena, California 91109, United States

<sup>2</sup>CSIRO Astronomy and Space Science, Canberra Deep Space Communication Complex – Australia

<sup>3</sup>European Space Agency – Germany

<sup>4</sup>JAXA – Japan

## Abstract

The X/Ka-band (8.4/32 GHz) Celestial Reference Frame became one component of the ICRF-3 in 2018. In the five years since, the X/Ka data set has increased by about a factor of two as well as adding the much needed north-south geometry from Japan to Australia. The latest solutions have median formal precisions of  $\sim 40 \mu\text{as}$  in  $\alpha \cos\delta$  and  $\sim 65 \mu\text{as}$  in  $\delta$ . The large spherical harmonics distortions seen in the ICRF3-XKa are greatly reduced with the Z-dipole term reduced from  $314 \mu\text{as}$  to statistical insignificance and with the quadrupole 2,0 magnetic term reduced to  $\sim 100 \mu\text{as}$ . Noting that the X/Ka frame is derived from a limited geometry of only five observing sites, we discuss the susceptibility of this frame to geometric distortion. Finally, we discuss the prospects for future improvements.

---

\*Speaker